## OoMA Fact Sheet

$G=(2 / 3) \times 10^{-7}$ dyne-cm ${ }^{2} /$ gram $^{2}$
$c=3 \times 10^{10} \mathrm{~cm} / \mathrm{sec}$
$k=(1 / 7) \times 10^{-15} \mathrm{erg} / \mathrm{K}$
$h=(2 / 3) \times 10^{-26}$ erg-sec
$\hbar=10^{-27}$ erg-sec
$N_{A}=6 \times 10^{23}$ nucleons/gram
$m_{p} / m_{e}=1836 \quad m_{p} c^{2}=938 \mathrm{MeV}$
$m_{e} \approx 10^{-27} \mathrm{gm} \quad m_{e} c^{2}=511 \mathrm{keV}$
$e=4.8 \times 10^{-10} \mathrm{esu}=1.6 \times 10^{-19}$ Coulomb
$\alpha=e^{2} / \hbar c=1 / 137$
$L_{\odot}=4 \times 10^{33} \mathrm{erg} / \mathrm{sec}$
Fusing H to He yields $0.7 \%$ of $m c^{2}$
He to $\mathrm{C} \& \mathrm{C}$ to Fe about $0.1 \%$ of $m c^{2}$ each
Solar Constant $=1.4 \mathrm{~kW} / \mathrm{m}^{2}$ at 1 AU
$M_{\odot}=2 \times 10^{33}$ grams $\quad R_{\odot}=7 \times 10^{10} \mathrm{~cm}$
$R_{\oplus}=6371 \mathrm{~km} \quad M_{\oplus}=3 \times 10^{-6} M_{\odot}$
$M_{J}=10^{-3} M_{\odot}$
Hubble radius $=c / H_{\circ}=1.3 \times 10^{28} \mathrm{~cm}$
Critical density $\sim 10^{-29} \mathrm{~g} / \mathrm{cm}^{3}$
$\sigma_{T}=(2 / 3) \times 10^{-24} \mathrm{~cm}^{2}$
$\sigma_{S B}=5.67 \times 10^{-5} \mathrm{erg} / \mathrm{cm}^{2} / \mathrm{sec} / \mathrm{K}^{4}$
Flux from a blackbody surface is $\sigma_{S B} T^{4}$
1 Farad $=9 \times 10^{11} \mathrm{~cm}$
1 ohm $=1 /\left(9 \times 10^{11}\right) \mathrm{sec} / \mathrm{cm}$
1 gram calorie $=4.2$ Watt-sec or Joules
Dietary calories are really kilocalories.
1 kiloton (kT) of TNT $=$ KE of 1000 met-
ric tonnes @ $2.9 \mathrm{~km} / \mathrm{sec} . \quad\left[1 \mathrm{kT}=10^{12}\right.$
gram-cal exactly]
Supernova kinetic energy $=10^{51}$ ergs
$1 \mathrm{AU}=(3 / 2) \times 10^{13} \mathrm{~cm}$
1 radian $=2 \times 10^{5}$ arc-seconds
1 square arcsec $=2.4 \times 10^{-11}$ steradians
$1 \mathrm{pc}=3 \times 10^{18} \mathrm{~cm}$
$1 \mathrm{erg}=6 \times 10^{11} \mathrm{eV}$
$1 \mathrm{eV} \sim 12,000 \mathrm{~K} \quad 1 \mathrm{eV} \sim 1.2 \mu \mathrm{~m}$
$h c / k \approx 1.44 \mathrm{~cm} \mathrm{~K}$
$1 \mathrm{Jy}=10^{-23} \mathrm{ergs} / \mathrm{cm}^{2} / \mathrm{sec} / \mathrm{Hz}$
1 year $\approx \pi \times 10^{7}$ seconds
1 Mpc is $1 \mathrm{~km} / \mathrm{sec}$ for 1000 Gyr
One atmosphere or 1 bar $=10^{6}$ dyne $/ \mathrm{cm}^{2}$
Maximum mass for white dwarfs: $1.4 \mathrm{M}_{\odot}$

Typical mass of neutron stars: $1.4 \mathrm{M}_{\odot}$

Stellar spectra - from "early" = hot to "late" = cool:
Oh Be A Fine Girl Kiss Me Later Tonight Luminosity class - the Roman numeral: "I" = supergiant = low surface gravity "III" = giant, "V" = dwarf = main sequence star $=$ high surface gravity.

| Sp.Type | $\log \left(L / L_{\odot}\right)$ | $\mathrm{M} / \mathrm{M}_{\odot}$ | $\mathrm{T}_{\text {eff }} \mathrm{K}$ |
| :--- | ---: | ---: | ---: |
| O5V | 5.82 | 40 | 40,000 |
| B0V | 4.66 | 18 | 28,000 |
| B5V | 2.94 | 9 | 15,500 |
| A0V | 1.78 | 3 | 9900 |
| A5V | 1.15 | 2 | 8500 |
| F0V | 0.88 | 1.7 | 7400 |
| F5V | 0.54 | 1.3 | 6580 |
| G0V | 0.15 | 1.1 | 6030 |
| G5V | -0.11 | 0.9 | 5520 |
| K0V | -0.38 | 0.8 | 4900 |
| K5V | -0.78 | 0.7 | 4130 |
| M0V | -1.22 | 0.5 | 3480 |
| M5V | -1.90 | 0.2 | 2800 |
| L0 | -3.65 |  | 2200 |
| L5 | -4.11 |  | 1700 |
| T0 | -4.57 |  | 1300 |
| T5 | -5.02 |  | 1000 |

1 magnitude is -4 db
A decibel ( db ) is a factor of $10^{0.1}$ in power.
$0^{\text {th }}$ mag at $\mathrm{V} \approx 10^{3}$ photons $/ \mathrm{cm}^{2} / \mathrm{sec} / \AA$.
$m_{b o l}=0$ for $2.5 \times 10^{-5} \mathrm{erg} / \mathrm{cm}^{2} / \mathrm{sec}$.
Bands central wavelengths in $\mu \mathrm{m}$ :
$\mathrm{U}=0.36, \mathrm{~B}=0.44, \mathrm{~V}=0.55, \mathrm{R}=0.7$, $\mathrm{I}=0.9, \mathrm{Z}=1.0, \mathrm{~J}=1.25, \mathrm{H}=1.6, \mathrm{~K}=$ $2.2, \mathrm{~L}=3.5, \mathrm{M}=4.6, \mathrm{~N}=10, \mathrm{Q}=20$
AB magnitudes have the same zeropoint flux in $F_{\nu}$ (3631 Jy) in all bands.
Johnson or "Vega" magnitudes have zeropoints that follow the spectrum of an A0V star.
$10^{n / 10}=1.26,1.6,2,2.5,3.2,4,5,6.3,8$.

