

UNIVERSITY OF KANSAS
 Department of Physics and Astronomy
 Physical Astronomy (ASTR 391) — Prof. Crossfield — Spring 2022

Problem Set 1

Due: Wednesdays, Feb 02, 2022, at the start of class (1100 Kansas Time)
 This problem set is worth **33 points**.

1. Astronomical Concepts [20 pts].

- (a) In a galaxy far, far away, the gas giant Endor orbits a Sun-like star at a distance of a_E . Endor (mass m_E) is orbited by a Forest Moon (m_m) with the same separation as found in the Earth-Moon system (a_D). What is the ratio (an algebraic expression, not just a number!) of the gravitational forces (i) between Endor and its star (mass m_*) and (ii) between Endor and its moon? Estimate which Force is stronger. [6 pts]

Solution: Since

$$F_G = G \frac{m_1 m_2}{r^2},$$

The ratio of the two forces is then

$$\frac{F_{E,*}}{F_{m,E}} = \frac{m_*}{m_m} \left(\frac{a_D}{1 \text{ au}} \right)^2.$$

I could look up a_D , or I could remember that light takes ~ 1 s to go from the Earth to the Moon and ~ 8 min from the Sun to the Earth (1 au). So the ratio of distances is $\sim 1/500$, and thus:

$$\boxed{\frac{F_{E,*}}{F_{m,E}} \approx \frac{m_*}{m_m} \frac{1}{2.5 \times 10^5}}.$$

Without knowing the mass m_m of the Forest Moon, we can't calculate an exact number. If it's like Earth's Moon, then $m_* \approx 3 \times 10^7 m_m$, but if it's a planet in its own right (say, twice the mass of the Earth) then $m_* \approx 1.5 \times 10^5$. So probably the star-planet force is stronger, but we can't be certain.

- (b) You have invented a matter-antimatter reactor that converts physical material into energy with 100% efficiency. Congratulations: you're a shoo-in for the Nobel Prize. (i) If you put 0.5 kg of matter (and an equal amount of antimatter) in your reactor, approximately how much energy (E_{reactor}) is released? (ii) If the reactor takes 0.5 s to use that fuel, what was its approximate power output in Solar Luminosities (L_\odot)? (iii) How does E_{reactor} compare to the total amount of energy used on Earth in a year? [7 pts]

Solution: We know $E = mc^2$, so the total energy released in this reaction is the mass energy of the fuel:

$$\begin{aligned} E &\approx (1 \text{ kg})(3 \times 10^8 \text{ m s}^{-1})^2 \\ &\approx \boxed{10^{17} \text{ J}}. \end{aligned}$$

When the energy is released in just 0.5 s, then the luminosity of your reactor is

$$L_{\text{reactor}} = \frac{\Delta E}{\Delta t} = \frac{10^{17} \text{ J}}{0.5 \text{ s}} = 4 \times 10^{17} \text{ W}.$$

Since $L_\odot \approx 4 \times 10^{26} \text{ W}$, we have

$$\boxed{L_{\text{reactor}} \approx 10^{-9} L_\odot}.$$

The world's annual energy consumption is roughly 500 million terajoules,

$$E_{\text{world}} \approx 500 \times 10^6 \times 10^{12} \text{ J} \approx 5 \times 10^{20} \text{ J}.$$

So your reactor produced an energy yield equal to about 0.0002 (0.02%) of the world's current energy output: still quite a feat!

- (c) Write the astronomer's version of the Ideal Gas Law. Explain each term (including its physical units), and how it might be used [7 pts].

Solution: The astrophysicist's ideal gas law is

$$P = nk_B T \quad (1)$$

where:

- i. P is the gas pressure (SI unit Pascals, $\text{Pa} = \text{N m}^{-2}$),
- ii. n is the gas number density (SI unit of inverse volume, m^{-3}) – can also be calculated if you know the gas density ($\rho = M/V = n\langle m \rangle$, where $\langle m \rangle$ is the average gas particle mass),
- iii. k_B is the Boltzmann constant, $\sim (1/7) \times 10^{-22} \text{ J/K}$, and
- iv. T is the temperature of the gas.

It is used to convert between the pressure, number density, and/or temperature of a gas in order to constrain any of those quantities that might not yet be known.

2. **Order-of-Magnitude Estimation [13 pts].** Strive to do as many of these calculations in your head (or with pencil and paper) as possible, aside from looking up any necessary physical constants.

- (a) **City on a Hill [5 pts.]** Roughly estimate the mass of Mount Oread, in kg and in M_\oplus (Earth masses).

Solution: From the topographic maps at <https://ngmdb.usgs.gov/topoview/> I estimate that Mt. Oread is about 2.5 km long, 700 m wide at the base, and 150 ft \approx 40 m higher than the surrounding land (Wikipedia tells me that is 60 m above downtown Lawrence, so your mileage may vary). I assume it has a triangular cross-sectional area, so

$$\begin{aligned} V_{\text{Oread}} &\approx \frac{1}{2} w \times h \times \ell \\ &\approx (350 \text{ m})(40 \text{ m})(2500 \text{ m}) \\ &\approx 3.5 \times 10^8 \text{ m}^3. \end{aligned}$$

Assuming a density the same as the Earth ($\sim 5000 \text{ kg m}^{-3}$), then we have

$$M_{\text{Oread}} \sim (5000)(3.5 \times 10^8) \text{ kg} \sim \boxed{1.8 \times 10^{12} \text{ kg}}. \quad (2)$$

Compared to the Earth's mass of $6 \times 10^{24} \text{ kg}$, KU's mountaintop perch is just $\sim \boxed{3 \times 10^{-13} M_\oplus}$.

A nice independent 'check' that a previous student pointed out is that the Great Pyramid of Giza, in Egypt, is estimated to weigh roughly six millions tons, $\approx 6 \times 10^9 \text{ kg}$. I've never been to Egypt, but it at least makes sense to me that Mt. Oread is bigger than a human-built pyramid: both have about the same height, but I estimated a base area much larger than that of the pyramid.

- (b) **How Big? [5 pts].** The French revolutionaries of the late 18th century defined the meter by setting the Earth's equator-to-pole distance to be 10,000 km. Estimate the radius (R_\oplus), volume (V_\oplus), and mass (M_\oplus) of the Earth, in SI units.

Solution: Since the Earth is approximately a sphere, its circumference $c_\oplus = 40,000 \text{ km} \approx 2\pi R_\oplus$. Thus,

$$R_\oplus \approx \frac{40,000 \text{ km}}{2\pi} \approx \frac{40,000 \text{ km}}{6} \boxed{6,500 \text{ km}} \quad (3)$$

which compares reasonably well with the modern value of $R_\oplus = 6,370 \text{ km}$.

Still approximating the Earth as a sphere, its volume is then just

$$V_\oplus = \frac{4}{3}\pi R_\oplus^3 \quad (4)$$

$$\approx 4(6,500 \text{ km})^3 \quad (5)$$

$$\approx 4 \times (6.5)^3 (10^6 \text{ m})^3 \quad (6)$$

$$\approx (26 \times 40) (10^{18} \text{ m}^3) \quad (7)$$

$$\approx (100) (10^{18} \text{ m}^3) \quad (8)$$

$$\approx \boxed{(10^{20} \text{ m}^3)} \quad (9)$$

$$(10)$$

... which compares well with the Earth's known volume of $V_{\oplus} = 1.083 \times 10^{20} \text{ m}^3$.

Mass is volume times density, i.e. $M_{\oplus} = V_{\oplus} \langle \rho_{\oplus} \rangle$. We know that water has $\rho_{\text{H}_2\text{O}} = 1.0 \text{ g/cc}$, and the Earth must be denser than this. We could look up the density of metals (7–10 g/cc) and assume that the Earth isn't solid metal and so must be *less* dense than these. (Of course, we could also just look up the average density of the Earth directly, but where's the fun in that?). If we assume $\langle \rho_{\oplus} \rangle \approx 4 \text{ g cm}^{-3} = 4000 \text{ kg m}^{-3}$, then we have

$$M_{\oplus} = V_{\oplus} \langle \rho_{\oplus} \rangle \tag{11}$$

$$\approx (10^{20} \text{ m}^3) (4000 \text{ kg m}^{-3}) \tag{12}$$

$$\boxed{\approx 4 \times 10^{24} \text{ kg m}^{-3}} \tag{13}$$

... which is a bit lower than the known value of $5.972 \times 10^{24} \text{ kg m}^{-3}$. We underestimated a bit because the Earth's actual density is 5.5 g/cc (not the 4.0 assumed above). But still: not too bad!

- (c) **How Big?!** [3 pts] Jupiter is roughly $10\times$ larger (in physical size) than the Earth (i.e., $R_{Jup} \approx 10R_{\oplus}$), and the Sun is roughly $10\times$ larger than Jupiter ($R_{\odot} \approx 10R_{Jup}$). Roughly estimate the volume of both of these objects, *relative to the volume of the Earth* (i.e., in units of V_{\oplus}).

Solution: Since all these objects are approximately spherical and $V \propto R^3$, we know

$$\frac{V_{Jup}}{V_{\oplus}} = \left(\frac{R_{Jup}}{R_{\oplus}} \right)^3 \approx 1000 \tag{14}$$

and so

$$\boxed{V_{Jup} \approx 1000V_{\oplus}}. \tag{15}$$

The actual value is 1322 or so, because actually Jupiter is a bit more than $10\times$ the size of the Earth.

For the Sun, we know

$$\frac{V_{\odot}}{V_{\oplus}} = \left(\frac{R_{\odot}}{R_{\oplus}} \right)^3 \approx 10^6 \tag{16}$$

and so

$$\boxed{V_{\odot} \approx 10^6V_{\oplus}}. \tag{17}$$

The actual value is around 1.31×10^6 : our estimate is off again because we just scaled from the Jovian value above, and Jupiter is actually bit more than $10\times$ the size of the Earth.